

## 10.2 [172 marks]

A planet has radius  $R$ . At a distance  $h$  above the surface of the planet the gravitational field strength is  $g$  and the gravitational potential is  $V$ .

- 1a. State what is meant by gravitational field strength.

[1 mark]

### Markscheme

the «gravitational» force per unit mass exerted on a point/small/test mass

[1 mark]

- 1b. Show that  $V = -g(R + h)$ .

[2 marks]

### Markscheme

at height  $h$  potential is  $V = -\frac{GM}{(R+h)}$

field is  $g = \frac{GM}{(R+h)^2}$

«dividing gives answer»

*Do not allow an answer that starts with  $g = -\frac{\Delta V}{\Delta r}$  and then cancels the deltas and substitutes  $R + h$*

[2 marks]

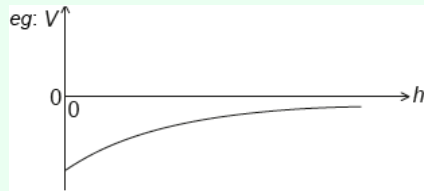
- 1c. Draw a graph, on the axes, to show the variation of the gravitational potential  $V$  of the planet with height  $h$  above the surface of the planet. [2 marks]



## Markscheme

correct shape and sign

non-zero negative vertical intercept



[2 marks]

- 1d. A planet has a radius of  $3.1 \times 10^6$  m. At a point P a distance  $2.4 \times 10^7$  m above the surface of the planet the gravitational field strength is  $2.2 \text{ N kg}^{-1}$ . Calculate the gravitational potential at point P, include an appropriate unit for your answer.

[1 mark]

## Markscheme

$$V = \left\langle -2.2 \times (3.1 \times 10^6 + 2.4 \times 10^7) \right\rangle \left\langle - \right\rangle 6.0 \times 10^7 \text{ J kg}^{-1}$$

Unit is essential

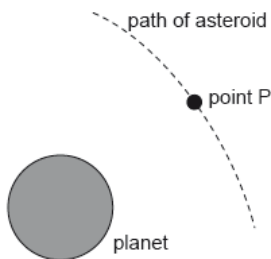
Allow eg  $\text{MJ kg}^{-1}$  if power of 10 is correct

Allow other correct SI units eg  $\text{m}^2\text{s}^{-2}$ ,  $\text{N m kg}^{-1}$

[1 mark]

- 1e. The diagram shows the path of an asteroid as it moves past the planet.

[3 marks]



When the asteroid was far away from the planet it had negligible speed. Estimate the speed of the asteroid at point P as defined in (b).

## Markscheme

total energy at P = 0 / KE gained = GPE lost

$$\left\langle \frac{1}{2}mv^2 + mV = 0 \Rightarrow v = \sqrt{-2V} \right\rangle$$

$$v = \left\langle \sqrt{2 \times 6.0 \times 10^7} \Rightarrow 1.1 \times 10^4 \text{ ms}^{-1} \right\rangle$$

Award [3] for a bald correct answer

Ignore negative sign errors in the workings

Allow ECF from 6(b)

[3 marks]

- 1f. The mass of the asteroid is  $6.2 \times 10^{12}$  kg. Calculate the gravitational force experienced by the **planet** when the asteroid is at point P. [2 marks]

## Markscheme

### ALTERNATIVE 1

force on asteroid is « $6.2 \times 10^{12} \times 2.2 \Rightarrow 1.4 \times 10^{13}$  «N»

«by Newton's third law» this is also the force on the planet

### ALTERNATIVE 2

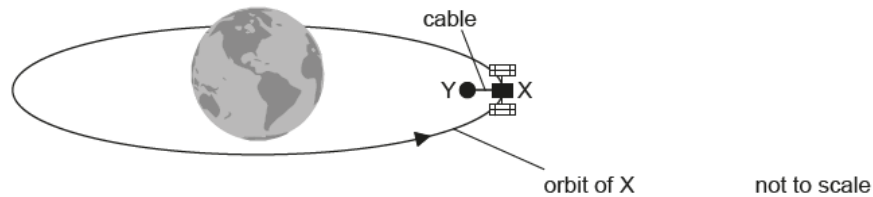
mass of planet =  $2.4 \times 10^{25}$  «kg» «from  $V = -\frac{GM}{(R+h)}$ »

force on planet «  
 $\frac{GMm}{(R+h)^2} = 1.4 \times 10^{13}$  «N»

MP2 must be explicit

[2 marks]

There is a proposal to power a space satellite X as it orbits the Earth. In this model, X is connected by an electronically-conducting cable to another smaller satellite Y.



- 2a. Satellite X orbits 6600 km from the centre of the Earth. [2 marks]  
 Mass of the Earth =  $6.0 \times 10^{24}$  kg  
 Show that the orbital speed of satellite X is about  $8 \text{ km s}^{-1}$ .

## Markscheme

$$\langle v = \sqrt{\frac{GME}{r}} \rangle = \sqrt{\frac{6.67 \times 10^{-11} \times 6.0 \times 10^{24}}{6600 \times 10^3}}$$

7800 «m s<sup>-1</sup>»

Full substitution required

Must see 2+ significant figures.

Satellite Y orbits closer to the centre of Earth than satellite X. Outline why

- 2b. the orbital times for X and Y are different. [1 mark]

## Markscheme

Y has smaller orbit/orbital speed is greater so time period is less

Allow answer from appropriate equation

Allow converse argument for X

- 2c. satellite Y requires a propulsion system. [2 marks]

## Markscheme

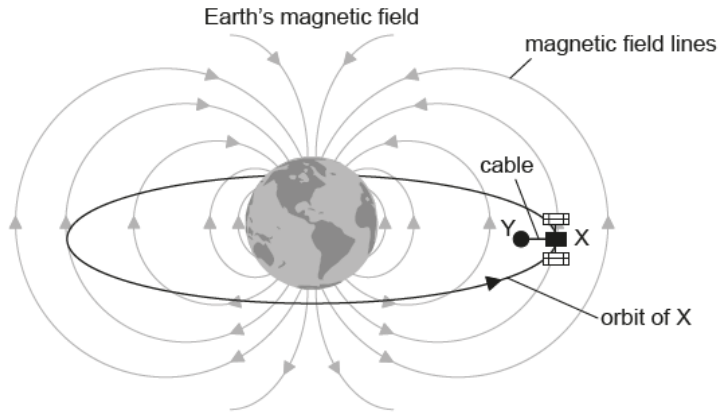
to stop Y from getting ahead

to remain stationary with respect to X

otherwise will add tension to cable/damage satellite/pull X out of its orbit

- 2d. The cable between the satellites cuts the magnetic field lines of the Earth at right angles.

[3 marks]



Explain why satellite X becomes positively charged.

## Markscheme

cable is a conductor and contains electrons

electrons/charges experience a force when moving in a magnetic field

use of a suitable hand rule to show that satellite Y becomes negative «so X becomes positive»

### **Alternative 2**

cable is a conductor

so current will flow by induction flow when it moves through a B field

use of a suitable hand rule to show current to right so «X becomes positive»

*Marks should be awarded from either one alternative or the other.*

*Do not allow discussion of positive charges moving towards X*

- 2e. Satellite X must release ions into the space between the satellites. Explain why the current in the cable will become zero unless there is a method for transferring charge from X to Y.

[3 marks]

## Markscheme

electrons would build up at satellite Y/positive charge at X

preventing further charge flow

by electrostatic repulsion

unless a complete circuit exists

- 2f. The magnetic field strength of the Earth is  $31 \mu\text{T}$  at the orbital radius of the satellites. The cable is 15 km in length. Calculate the emf induced in the cable.

[2 marks]

## Markscheme

$$\ll \mathcal{E} = Blv \Rightarrow 31 \times 10^{-6} \times 7990 \times 15000$$

$$3600 \ll V \gg$$

Allow 3700 «V» from  $v = 8000 \text{ m s}^{-1}$ .

The cable acts as a spring. Satellite Y has a mass  $m$  of  $3.5 \times 10^2 \text{ kg}$ . Under certain circumstances, satellite Y will perform simple harmonic motion (SHM) with a period  $T$  of 5.2 s.

- 2g. Estimate the value of  $k$  in the following expression.

[3 marks]

$$T = 2\pi\sqrt{\frac{m}{k}}$$

Give an appropriate unit for your answer. Ignore the mass of the cable and any oscillation of satellite X.

## Markscheme

$$\text{use of } k = \ll \frac{4\pi^2 m}{T^2} \Rightarrow \frac{4 \times \pi^2 \times 350}{5.2^2}$$

510

$\text{N m}^{-1}$  **or**  $\text{kg s}^{-2}$

Allow MP1 and MP2 for a bald correct answer

Allow 500

Allow N/m etc.

- 2h. Describe the energy changes in the satellite Y-cable system during one cycle of the oscillation.

[2 marks]

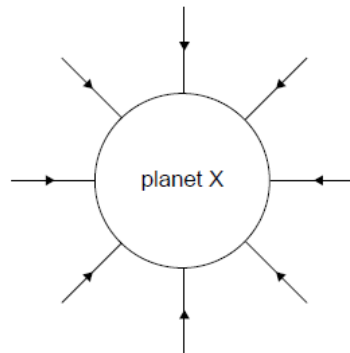
## Markscheme

$E_p$  in the cable/system transfers to  $E_k$  of Y

and back again twice in each cycle

Exclusive use of gravitational potential energy negates MP1

The diagram shows the gravitational field lines of planet X.



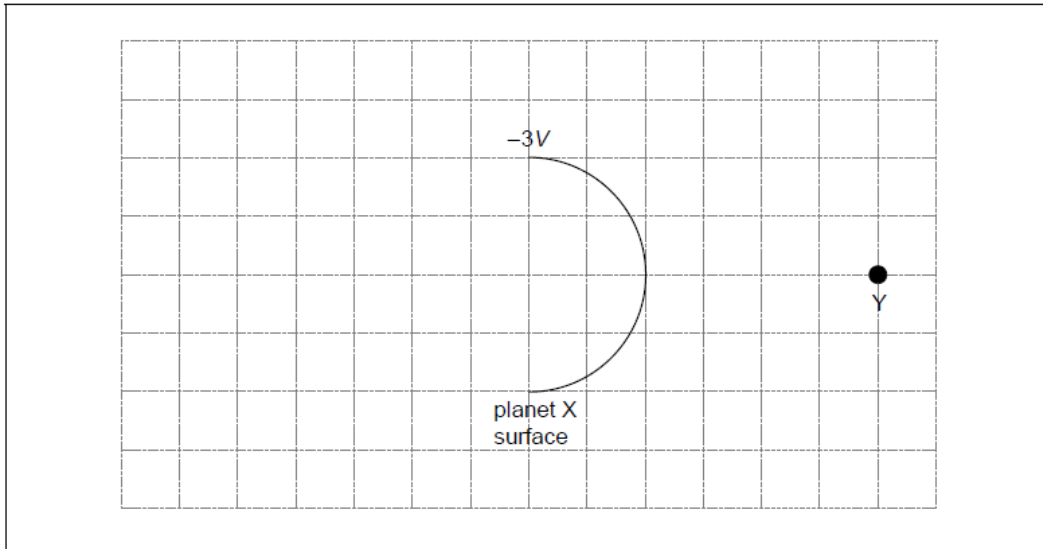
- 3a. Outline how this diagram shows that the gravitational field strength of planet X decreases with distance from the surface.

[1 mark]

## Markscheme

the field lines/arrows are further apart at greater distances from the surface

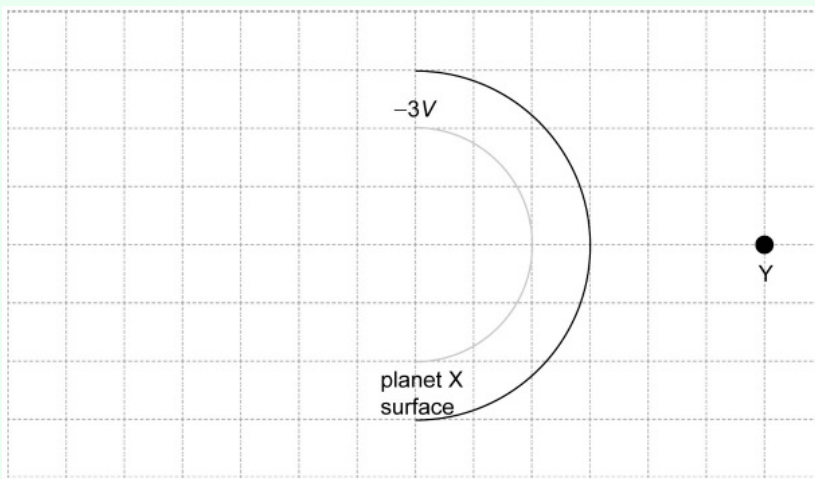
- 3b. The diagram shows part of the surface of planet X. The gravitational potential at the surface of planet X is  $-3 V$  and the gravitational potential at point Y is  $-V$ . [2 marks]



Sketch on the grid the equipotential surface corresponding to a gravitational potential of  $-2 V$ .

## Markscheme

circle centred on Planet X  
three units from Planet X centre



- 3c. A meteorite, very far from planet X begins to fall to the surface with a negligibly small initial speed. The mass of planet X is  $3.1 \times 10^{21}$  kg and its radius is  $1.2 \times 10^6$  m. The planet has no atmosphere. Calculate the speed at which the meteorite will hit the surface. [3 marks]

## Markscheme

$$\text{loss in gravitational potential} = \frac{6.67 \times 10^{-11} \times 3.1 \times 10^{21}}{1.2 \times 10^6}$$

$$\ll = 1.72 \times 10^5 \text{ J Kg}^{-1} \gg$$

$$\text{equate to } \frac{1}{2}v^2$$

$$v = 590 \text{ «m s}^{-1}\gg$$

Allow ECF from MP1.

- 3d. At the instant of impact the meteorite which is made of ice has a temperature of 0 °C. Assume that all the kinetic energy at impact gets transferred into internal energy in the meteorite. Calculate the percentage of the meteorite's mass that melts. The specific latent heat of fusion of ice is  $3.3 \times 10^5 \text{ J kg}^{-1}$ . [2 marks]

## Markscheme

available energy to melt one kg  $1.72 \times 10^5 \text{ «J»}$

fraction that melts is  $\frac{1.72 \times 10^5}{3.3 \times 10^5} = 0.52$  **OR** 52%

Allow ECF from MP1.

Allow 53% from use of  $590 \text{ ms}^{-1}$ .

The gravitational potential due to the Sun at its surface is  $-1.9 \times 10^{11} \text{ J kg}^{-1}$ . The following data are available.

Mass of Earth	= $6.0 \times 10^{24} \text{ kg}$
Distance from Earth to Sun	= $1.5 \times 10^{11} \text{ m}$
Radius of Sun	= $7.0 \times 10^8 \text{ m}$

- 4a. Outline why the gravitational potential is negative. [2 marks]

## Markscheme

potential is defined to be zero at infinity

so a positive amount of work needs to be supplied for a mass to reach infinity

- 4b. The gravitational potential due to the Sun at a distance  $r$  from its centre is  $V_S$ . Show that  $rV_S = \text{constant}$ . [1 mark]

## Markscheme

$V_S = -\frac{GM}{r}$  so  $r \times V_S = -GM = \text{constant}$  because  $G$  and  $M$  are constants

- 4c. Calculate the gravitational potential energy of the Earth in its orbit around the Sun. Give your answer to an appropriate number of significant figures. [2 marks]

## Markscheme

$GM = 1.33 \times 10^{20} \text{ «J m kg}^{-1}\text{»}$

GPE at Earth orbit  $\text{«} = -\frac{1.33 \times 10^{20} \times 6.0 \times 10^{24}}{1.5 \times 10^{11}} = \text{«} \rightarrow 5.3 \times 10^{33} \text{ «J»}$

Award [1 max] unless answer is to 2 sf.

Ignore addition of Sun radius to radius of Earth orbit.

- 4d. Calculate the total energy of the Earth in its orbit. [2 marks]

## Markscheme

### ALTERNATIVE 1

work leading to statement that kinetic energy  $\frac{GMm}{2r}$ , **AND** kinetic energy evaluated to be «+»  $2.7 \times 10^{33}$  «J»

energy «= PE + KE = answer to (b)(ii) +  $2.7 \times 10^{33}$ » = «-»  $2.7 \times 10^{33}$  «J»

### ALTERNATIVE 2

statement that kinetic energy is  $= -\frac{1}{2}$  gravitational potential energy in orbit

so energy «=  $\frac{\text{answer to (b)(ii)}}{2}$ » = «-»  $2.7 \times 10^{33}$  «J»

*Various approaches possible.*

- 4e. An asteroid strikes the Earth and causes the orbital speed of the Earth to suddenly decrease. Suggest the ways in which the orbit of the Earth will change. [2 marks]

## Markscheme

«KE will initially decrease so» total energy decreases

**OR**

«KE will initially decrease so» total energy becomes more negative

Earth moves closer to Sun

new orbit with greater speed «but lower total energy»

changes ellipticity of orbit

- 4f. Outline, in terms of the force acting on it, why the Earth remains in a circular orbit around the Sun. [2 marks]

## Markscheme

centripetal force is required

and is provided by gravitational force between Earth and Sun

*Award [1 max] for statement that there is a "centripetal force of gravity" without further qualification.*

- 5a. Explain what is meant by the gravitational potential at the surface of a planet. [2 marks]

## Markscheme

the «gravitational» work done «by an external agent» per/on unit mass/kg

*Allow definition in terms of reverse process of moving mass to infinity eg "work done on external agent by..."*

*Allow "energy" as equivalent to "work done"*

in moving a «small» mass from infinity to the «surface of» planet / to a point

**N.B.:** on SL paper Q5(a)(i) and (ii) is about "gravitational field".



- 5b. An unpowered projectile is fired vertically upwards into deep space from the surface of planet Venus. Assume that the gravitational effects of the Sun and the other planets are negligible. [5 marks]

The following data are available.

Mass of Venus =  $4.87 \times 10^{24}$  kg Radius of Venus =  $6.05 \times 10^6$  m Mass of projectile =  $3.50 \times 10^3$  kg Initial speed of projectile =  $1.10 \times$  escape speed

- (i) Determine the initial kinetic energy of the projectile.  
(ii) Describe the subsequent motion of the projectile until it is effectively beyond the gravitational field of Venus.

## Markscheme

i

escape speed

Care with ECF from MP1.

$$v = \sqrt{\left(\frac{2GM}{R}\right)} \Rightarrow$$

$$\sqrt{\left(\frac{2 \times 6.67 \times 10^{-11} \times 4.87 \times 10^{24}}{6.05 \times 10^6}\right)} \text{ or } 1.04 \times 10^4 \text{ «m s}^{-1}\text{»}$$

$$\text{or } \ll 1.1 \times 1.04 \times 10^4 \text{ m s}^{-1} \gg = 1.14 \times 10^4 \text{ «m s}^{-1}\text{»}$$

$$\text{KE} = \ll 0.5 \times 3500 \times (1.1 \times 1.04 \times 10^4 \text{ m s}^{-1})^2 \gg = 2.27 \times 10^{11} \text{ «J»}$$

Award [1 max] for omission of 1.1 – leads to  $1.88 \times 10^{11} \text{ m s}^{-1}$ .

Award [2] for a bald correct answer.

ii

Velocity/speed decreases / projectile slows down «at decreasing rate»

«magnitude of» deceleration decreases «at decreasing rate»

Mention of deceleration scores MP1 automatically.

velocity becomes constant/non-zero

**OR**

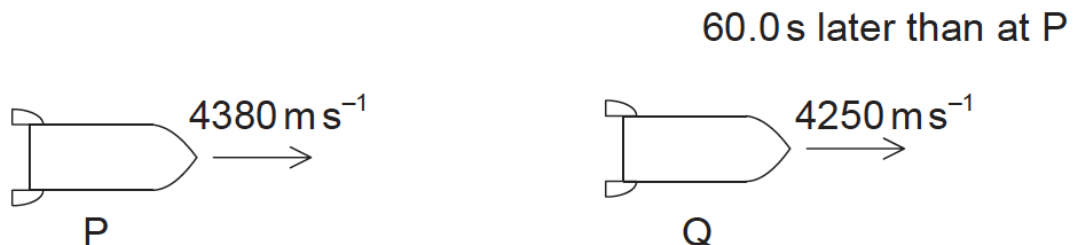
deceleration tends to zero

Accept “negative acceleration” for “deceleration”.

Must see “velocity” not “speed” for MP3.

### Part 2 Motion of a rocket

A rocket is moving away from a planet within the gravitational field of the planet. When the rocket is at position P a distance of  $1.30 \times 10^7$  m from the centre of the planet, the engine is switched off. At P, the speed of the rocket is  $4.38 \times 10^3 \text{ ms}^{-1}$ .



At a time of 60.0 s later, the rocket has reached position Q. The speed of the rocket at Q is  $4.25 \times 10^3 \text{ ms}^{-1}$ . Air resistance is negligible.

- 6a. Outline, with reference to the energy of the rocket, why the speed of the rocket is changing between P and Q. [2 marks]

## Markscheme

gravitational potential energy is being gained;  
this is at the expense of kinetic energy (and speed falls);

6b. Estimate the average gravitational field strength of the planet between P and Q.

[2 marks]

## Markscheme

$$\left( \text{acceleration} = \frac{(v-u)}{t} = \frac{4.25 \times 10^3 - 4.38 \times 10^3}{60} \right) (-) 2.17 \text{ (ms}^{-2}\text{)};$$

gravitational field strength = acceleration of rocket (=2.17 N kg<sup>-1</sup>); } (allow  $g = a$  in symbols)

**or**

computes potential difference from KE per unit mass change (5.61 × 10<sup>5</sup>),

computes distance travelled (0.259 Mm), uses

$$g = \frac{(-)\Delta V}{\Delta r};$$

$$g = (-) 2.17 \text{ (ms}^{-2}\text{)};$$

6c. A space station is in orbit at a distance  $r$  from the centre of the planet in (e)(i). A satellite is launched from the space station so as just to escape from the gravitational field of the planet. The launch takes place in the same direction as the velocity of the space station. Outline why the launch velocity relative to the space station can be less than your answer to (e)(i).

[1 mark]

## Markscheme

the satellite has velocity/kinetic energy as it is orbiting with the space station;

This question is about the energy of an orbiting satellite.

A space shuttle of mass  $m$  is launched in the direction of the Earth's South Pole.

7a. The kinetic energy  $E_K$  given to the shuttle at its launch is given by the expression

[2 marks]

$$E_K = \frac{7GMm}{8R_E}$$

where  $G$  is the gravitational constant,  $M$  is mass of the Earth and  $R_E$  is the radius of the Earth. Deduce that the shuttle cannot escape the gravitational field of the Earth.

## Markscheme

KE needs to be  $\geq$  (magnitude of) GPE at surface  $\left( -\frac{GMm}{R_E} \right)$ ;

But KE is  $\frac{7GMm}{8R_E} < \frac{GMm}{R_E}$  / OWTTE;

**or**

shows that total energy at launch =  $-\frac{GMm}{8R_E}$ ; (appropriate working required)

this is  $< 0$ , so escape impossible;

**or**

states that escape velocity needed is  $\sqrt{\frac{2GM}{R_E}}$ ;

shows launch velocity is only  $\sqrt{\frac{7GM}{4R_E}}$ ; (appropriate working required)

The shuttle enters a circular orbit of radius  $R$  around the Earth.

7b. Show that the total energy of the shuttle in its orbit is given by  $-\frac{GMm}{2R}$ . Air resistance is negligible.

[3 marks]

## Markscheme

$$E_{\text{tot}} = \text{PE} + \text{KE};$$

shows that kinetic energy =  $(\frac{1}{2}mv^2 =) \frac{GMm}{2R}$ ; (appropriate working required)

adds PE  $(-\frac{GMm}{R})$  and KE to get given answer; (appropriate working required)

- 7c. Using the expression for  $E_K$  in (a) and your answer to (b)(i), determine  $R$  in terms of  $R_E$ .

[3 marks]

## Markscheme

$$-\frac{GMm}{R_E} + \frac{7GMm}{8R_E} = -\frac{GMm}{2R}; \text{ (equating total energy at launch and in orbit)}$$

$$\frac{1}{8R_E} = \frac{1}{2R};$$

$$R = 4R_E;$$

Award [0] for an answer such as  $R = \frac{4R_E}{7}$ .

- 7d. In practice, the total energy of the shuttle decreases as it collides with air molecules in the upper atmosphere. Outline what happens to the speed of the shuttle when this occurs.

[2 marks]

## Markscheme

total energy decreases/becomes a greater negative value, so  $R$  decreases;

as  $R$  decreases kinetic energy increases;

speed increases;

Allow third marking point even if reasoning is incorrect.

This question is in **two** parts. **Part 1** is about the use of renewable energy sources. **Part 2** is about the gravitational potential of the Earth.

**Part 2** Gravitational potential of the Earth

8. The table gives the gravitational potential  $V$  for various distances  $r$  from the **surface** of Earth. The radius of Earth is  $6.4 \times 10^3$  km. [3 marks]

$V / 10^7 \text{ J kg}^{-1}$	$r / 10^3 \text{ km}$
-6.24	0
-3.84	4.0
-1.04	32

Show that the data are consistent with Earth acting as a point mass with its mass concentrated at its centre.

## Markscheme

recognize that  $V \times$  a distance is constant /  $\frac{V_1}{V_2} = \frac{R_2}{R_1} (= 1.625)$ ;

use of  $R_E + r$ ;

evaluates at least two data points  $4.0 \times 10^{14}$ ; (allow  $V \times R$  to be expressed in any consistent unit)

Award [2 max] if answer fails to use radius of Earth but infers wrong conclusion from correct evaluation of two or more data points.

Award [3 max] if answer evaluates mass of Earth and shows that it is the same for two or more data points.

Part 2 Gravitational potential

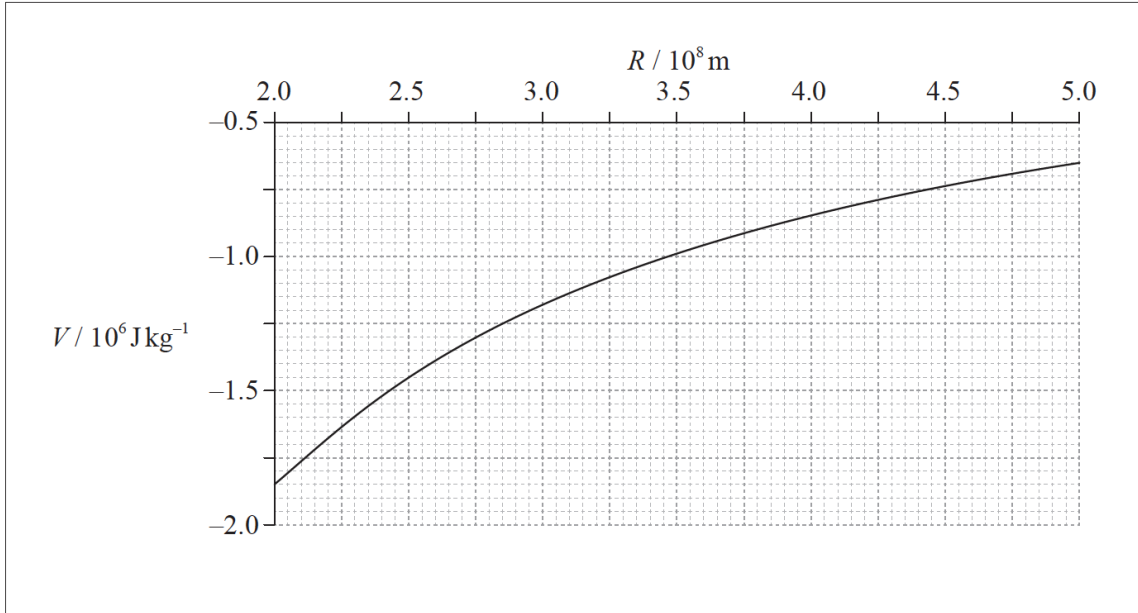
9a. Define *gravitational potential* at a point in a gravitational field.

[3 marks]

## Markscheme

work done per unit mass;  
in bringing (test) mass from infinity to point;  
reference to small/point (test) mass;

9b. The graph shows how the gravitational potential  $V$  of Earth varies with distance  $R$  from the centre of Earth in the range  $R=2.0 \times 10^8$  m to  $R=5.0 \times 10^8$  m. [6 marks]



The Moon is at a distance of  $4.0 \times 10^8$  m from the centre of Earth. At some time in the past it was at a distance of  $2.7 \times 10^8$  m from the centre of Earth.

Use the graph opposite to determine

- the present day magnitude of the acceleration of the Moon.
- by how much the potential energy of the Moon has changed as a result of moving from  $R=2.7 \times 10^8$  m to  $R=4.0 \times 10^8$  m. The mass of the Moon is  $7.4 \times 10^{22}$  kg.

## Markscheme

(i) tangent construction attempted at  $R=4.0 \times 10^8$  m;  
triangle/pair of coordinates used in calculation;  
attempt to calculate gradient;  
 $2.5 \times 10^{-3} \text{ ms}^{-2}$ ; (accept answers in the range of 2.2 to 2.7)

Award [1 max] for  
 $\frac{V}{R}$  to give  $(-2.1 \times 10^{-3})$ .

(ii) change in  $V=0.45 \times 10^6 - 0.50 \times 10^6 \text{ J kg}^{-1}$ ;  
change in PE =  $(0.5 \times 10^6 \times 7.4 \times 10^{22}) - 3.3 - 3.7 \times 10^{28} \text{ J}$ ;

9c. State why the change of potential energy in (f)(ii) is an increase.

[1 mark]

## Markscheme

work is done against the gravitational field of Earth / Moon is now closer to infinity/further from Earth /  $\frac{-GMm}{R}$  means that as R increases potential increases/becomes less negative;

This question is in two parts. **Part 1** is about electric charge and electric circuits. **Part 2** is about momentum.

**Part 1** Electric charge and electric circuits

10a. State Coulomb's law.

[2 marks]

## Markscheme

the force between two (point) charges;

is inversely proportional to the square of their separation and (directly) proportional to (the product of) their magnitudes;

Allow [2] for equation with  $F$ ,  $Q$  and  $r$  defined.

10b. In a simple model of the hydrogen atom, the electron can be regarded as being in a circular orbit about the proton. The radius of the orbit is  $2.0 \times 10^{-10}$  m. [7 marks]

(i) Determine the magnitude of the electric force between the proton and the electron.

(ii) Calculate the magnitude of the electric field strength  $E$  and state the direction of the electric field due to the proton at a distance of  $2.0 \times 10^{-10}$  m from the proton.

(iii) The magnitude of the gravitational field due to the proton at a distance of  $2.0 \times 10^{-10}$  m from the proton is  $H$ . Show that the ratio  $\frac{H}{E}$  is of the order  $10^{-28} \text{C kg}^{-1}$ .

(iv) The orbital electron is transferred from its orbit to a point where the potential is zero. The gain in potential energy of the electron is  $5.4 \times 10^{-19} \text{J}$ . Calculate the value of the potential difference through which the electron is moved.

## Markscheme

$$(i) F = \left( k \frac{q_1 q_2}{r^2} = \right) \frac{9 \times 10^9 \times [1.6 \times 10^{-19}]^2}{4 \times 10^{-20}};$$

$$= 5.8 \times 10^{-9} (\text{N});$$

Award [0] for use of masses in place of charges.

$$(ii) \left( \frac{(b)(i)}{1.6 \times 10^{-19}} \text{ or } 3.6 \times 10^{10} (\text{NC}^{-1}) \text{ or } (\text{Vm}^{-1}); \right)$$

(directed) away from the proton;

Allow ECF from (b)(i).

Do not penalize use of masses in both (b)(i) and (b)(ii) – allow ECF.

$$(iii) H = \left( G \frac{m}{r^2} = \right) \frac{6.67 \times 10^{-11} \times 1.673 \times 10^{-27}}{4 \times 10^{-20}} = 2.8 \times 10^{-18} (\text{Nkg}^{-1});$$

$$\frac{H}{E} = \frac{2.8 \times 10^{-18}}{3.6 \times 10^{10}} \text{ or } 7.8 \times 10^{-29} (\text{Ckg}^{-1});$$

$$(\approx 10^{28} \text{Ckg}^{-1})$$

Allow ECF from (b)(i).

$$(iv) 3.4 (\text{V});$$

10c. An electric cell is a device that is used to transfer energy to electrons in a circuit. A particular circuit consists of a cell of emf  $\epsilon$  and internal resistance  $r$  connected in series with a resistor of resistance  $5.0 \Omega$ . [6 marks]

(i) Define *emf of a cell*.

(ii) The energy supplied by the cell to one electron in transferring it around the circuit is  $5.1 \times 10^{-19} \text{J}$ . Show that the emf of the cell is  $3.2 \text{V}$ .

(iii) Each electron in the circuit transfers an energy of  $4.0 \times 10^{-19} \text{J}$  to the  $5.0 \Omega$  resistor. Determine the value of the internal resistance  $r$ .

## Markscheme

(i) power supplied per unit current / energy supplied per unit charge / work done per unit charge;

(ii) energy supplied per coulomb =  $\frac{5.1 \times 10^{-19}}{1.6 \times 10^{-19}}$  or 3.19(V);  
( $\approx 3.2$ V)

(iii) pd across  $5.0\Omega$  resistor =  $\left(\frac{4.0 \times 10^{-19}}{1.6 \times 10^{-19}}\right) = 2.5$  (V);  
pd across  $r = (3.2 - 2.5) = 0.70$  (V);

and

**either**

current in circuit =  $\left(\frac{2.5}{5.0}\right) = 0.5$  (A);  
resistance of  $r = \left(\frac{0.70}{0.50}\right) = 1.4$  ( $\Omega$ );

or

resistance of  $r = \frac{0.70}{2.5} \times 5.0$ ;  
 $= 1.4$  ( $\Omega$ );

or

$3.2 = 0.5(R+r)$ ;  
resistance of  $r = 1.4$  ( $\Omega$ );  
Award [4] for alternative working leading to correct answer.  
Award [4] for a bald correct answer.

This question is in **two** parts. **Part 1** is about gravitational force fields. **Part 2** is about properties of a gas.

**Part 1** Gravitational force fields

11a. State Newton's universal law of gravitation.

[2 marks]

## Markscheme

the (attractive) force between two (point) masses is directly proportional to the product of the masses;  
and inversely proportional to the square of the distance (between their centres of mass);

Use of equation is acceptable:

Award [2] if all five quantities defined. Award [1] if four quantities defined.

11b. A satellite of mass  $m$  orbits a planet of mass  $M$ . Derive the following relationship between the period of the satellite  $T$  and the radius of its orbit  $R$  (Kepler's third law).

[3 marks]

$$T^2 = \frac{4\pi^2 R^3}{GM}$$

## Markscheme

$$G \frac{Mm}{R^2} = \frac{mv^2}{R} \text{ so } v^2 = \frac{GM}{R};$$

$$v = \frac{2\pi R}{T};$$

$$v^2 = \frac{4\pi^2 R^2}{T^2} = \frac{GM}{R};$$

or

$$G \frac{Mm}{R^2} = m\omega^2 R;$$

$$\omega^2 = \frac{4\pi^2}{T^2};$$

$$\frac{4\pi^2}{T^2} = \frac{GM}{R^3};$$

Award [3] to a clear response with a missing step.

- 11c. A polar orbiting satellite has an orbit which passes above both of the Earth's poles. One polar orbiting satellite used for Earth observation has an orbital period of  $6.00 \times 10^3$ s. [8 marks]

$$\begin{aligned} \text{Mass of Earth} &= 5.97 \times 10^{24} \text{ kg} \\ \text{Average radius of Earth} &= 6.37 \times 10^6 \text{ m} \end{aligned}$$

- (i) Using the relationship in (b), show that the average height above the surface of the Earth for this satellite is about 800 km.  
 (ii) The satellite moves from an orbit of radius 1200 km above the Earth to one of radius 2500 km. The mass of the satellite is 45 kg.  
 Calculate the change in the gravitational potential energy of the satellite.

- (iii) Explain whether the gravitational potential energy has increased, decreased or stayed the same when the orbit changes, as in (c)(ii).

## Markscheme

$$(i) R^3 = \frac{6.67 \times 10^{-11} \times 5.97 \times 10^{24} \times 6000^2}{4 \times \pi^2};$$

$$R = 7.13 \times 10^6 \text{ (m)};$$

$$h = (7.13 \times 10^6 - 6.37 \times 10^6) = 760 \text{ (km)};$$

Award [3] for an answer of 740 with  $\pi$  taken as 3.14.

$$(ii) \text{ clear use of } \Delta V = \frac{\Delta E}{m} \text{ and } V = -\frac{Gm}{r} \text{ or } \Delta E = GMm \left( \frac{1}{r_1} - \frac{1}{r_2} \right);$$

one value of potential energy calculated ( $2.37 \times 10^9$  or  $2.02 \times 10^9$ );

$$3.5 \times 10^8 \text{ (J)};$$

Award [3] for a bald correct answer.

Award [2] for  $7.7 \times 10^9$ . Award [1] for  $7.7 \times 10^{12}$ .

Award [0] for answers using  $mg\Delta h$ .

(iii) increased;

further from Earth / closer to infinity / smaller negative value;

Award [0] for a bald correct answer.

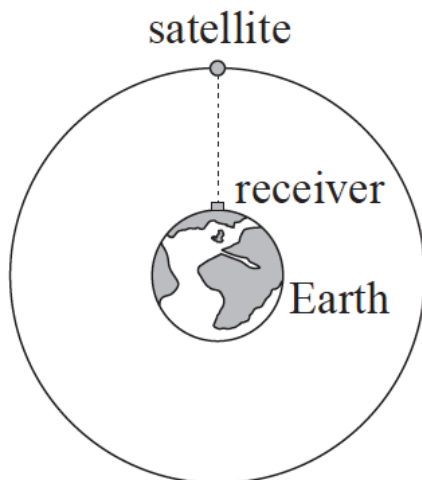
### Part 2 Satellite

- 12a. State, in words, Newton's universal law of gravitation. [2 marks]

## Markscheme

force is proportional to product of masses and inversely proportional to square of distance apart;  
 reference to point masses;

- 12b. The diagram shows a satellite orbiting the Earth. The satellite is part of the network of global-positioning satellites (GPS) that transmit radio signals used to locate the position of receivers that are located on the Earth. [3 marks]



(not to scale)

When the satellite is directly overhead, the microwave signal reaches the receiver 67ms after it leaves the satellite.

- (i) State the order of magnitude of the wavelength of microwaves.  
 (ii) Calculate the height of the satellite above the surface of the Earth

## Markscheme

- (i) order of 1 cm;  
 (ii)  $3 \times 10^8 \times 67 \times 10^{-3}$ ;  
 $2.0 \times 10^7 \text{m}$ ;

- 12c. (i) Explain why the satellite is accelerating towards the centre of the Earth even though its orbital speed is constant. [8 marks]  
 (ii) Calculate the gravitational field strength due to the Earth at the position of the satellite.  
 Mass of Earth =  $6.0 \times 10^{24} \text{kg}$   
 Radius of Earth =  $6.4 \times 10^6 \text{m}$   
 (iii) Determine the orbital speed of the satellite.  
 (iv) Determine, in hours, the orbital period of the satellite.

## Markscheme

- (i) force required towards centre of Earth to maintain orbit;  
 force means that there is an acceleration / *OWTTE*;

**or**

direction changes;  
 a change in velocity therefore acceleration;

(ii) uses =  $\frac{GM}{r^2}$  **or**  $\frac{6.7 \times 10^{-11} \times 6.0 \times 10^{24}}{[2.6 \times 10^7]^2}$ ;  
 $0.57 \text{Nkg}^{-1}$ ; (allow  $\text{ms}^{-2}$ )

(iii)  $v = \sqrt{0.57 \times (2.0 \times 10^7 + 6.4 \times 10^6)}$  by equating  $\frac{v^2}{r}$  and  $g$ ;  
 $3900 \text{ms}^{-1}$ ;

(iv)  $T = 2\pi \frac{2.6 \times 10^7}{3900}$ ;  
 11.9 hours;

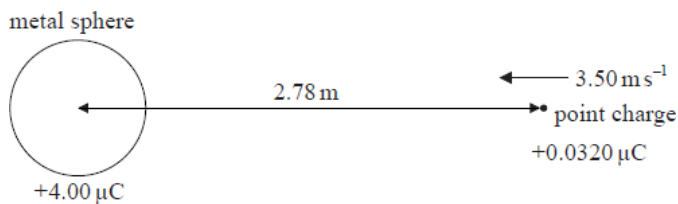
This question is about electric potential.

- 13a. Define *electric potential* at a point in an electric field. [3 marks]

## Markscheme

the work done per unit charge;  
 when a small/test/point positive charge; (*charge sign is essential*)  
 is moved from infinity to the point;

- 13b. A positive point charge is moving towards a small, charged metal sphere along a radial path. [6 marks]



At the position shown in the diagram, the point charge has a speed of  $3.50 \text{ m s}^{-1}$  and is at a distance of  $2.78 \text{ m}$  from the centre of the metal sphere. The charge on the sphere is  $+4.00 \mu\text{C}$ .

- (i) State the direction of the velocity of the point charge with respect to an equipotential surface due to the metal sphere.  
 (ii) Show that the electric potential  $V$  due to the charged sphere at a distance of  $2.78 \text{ m}$  from its centre is  $1.29 \times 10^4 \text{ V}$ .  
 (iii) The electric potential at the surface of the sphere is  $7.20 \times 10^4 \text{ V}$ . The point charge has a charge of  $+0.0320 \mu\text{C}$  and its mass is  $1.20 \times 10^{-4} \text{ kg}$ . Determine if the point charge will collide with the metal sphere.



## Markscheme

(i) perpendicular / at right angles / at  $90^\circ$  / normal;

(ii)  $V = \frac{8.99 \times 10^9 \times 4.00 \times 10^{-6}}{2.78}$  or  $1.2935 \times 10^4 \text{V}$ ; (use of  $\frac{1}{4\pi\epsilon_0}$  gives  $1.29378 \times 10^4$ )  
( $\approx 1.29 \times 10^4 \text{V}$ )

(iii) difference in potential =  $(7.20 \times 10^4 - 1.29 \times 10^4) = 5.91 \times 10^4$ ;

required loss in kinetic energy/minimum kinetic energy to reach sphere =  $(0.032 \times 10^{-6} \times 5.91 \times 10^4) = 1.89 \times 10^{-3} \text{J}$ ;

available kinetic energy =  $(\frac{1}{2} \times 1.20 \times 10^{-4} \times 3.50^2) = 7.35 \times 10^{-4} \text{J}$ ; not enough (initial) kinetic energy to reach sphere;

Response needs some statement of conclusion, e.g. so it does not reach sphere.

Allow answer in terms of minimum speed to reach sphere  $5.61 \text{ms}^{-1}$ .

This question is about escape speed and gravitational effects.

14a. Explain what is meant by escape speed.

[2 marks]

## Markscheme

(minimum) speed of object to escape gravitational field of a planet/travel to infinity;

at surface of planet;

without (further) energy input;

14b. Titania is a moon that orbits the planet Uranus. The mass of Titania is  $3.5 \times 10^{21} \text{kg}$ . The radius of Titania is 800 km.

[5 marks]

(i) Use the data to calculate the gravitational potential at the surface of Titania.

(ii) Use your answer to (b)(i) to determine the escape speed for Titania.

## Markscheme

(i)  $-\frac{6.67 \times 10^{-11} \times 3.5 \times 10^{21}}{8.0 \times 10^5}$ ;  
 $-2.9 \times 10^5 \text{Jkg}^{-1}$ ; (allow  $\text{Nmkg}^{-1}$ )

Award [1 max] if negative sign omitted.

(ii)  $\frac{1}{2}mv^2 = mV$ ;

speed =  $\sqrt{2 \times 2.9 \times 10^5}$ ; (allow ECF from (b)(i))

$7.6 \times 10^2 \text{ms}^{-1}$ ;

Ignore sign.

Award [3] for a bald correct answer.

14c. An astronaut visiting Titania throws an object away from him with an initial horizontal velocity of  $1.8 \text{ m s}^{-1}$ . The object is 1.5 m above the moon's surface when it is thrown. The gravitational field strength at the surface of Titania is  $0.37 \text{ N kg}^{-1}$ .

[3 marks]

Calculate the distance from the astronaut at which the object first strikes the surface.

## Markscheme

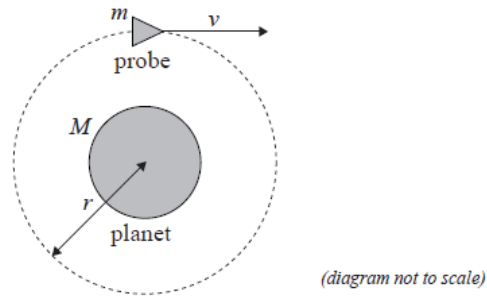
time to hit surface =  $\sqrt{\frac{2.0 \times 1.5}{0.37}}$  (= 2.85s);

distance to impact =  $2.85 \times 1.8$ ;

5.1m;

This question is about a probe in orbit.

A probe of mass  $m$  is in a circular orbit of radius  $r$  around a spherical planet of mass  $M$ .



- 15a. State why the work done by the gravitational force during one full revolution of the probe is zero.

[1 marks]

## Markscheme

because the force is always at right angles to the velocity / motion/orbit is an equipotential surface;  
*Do not accept answers based on the displacement being zero for a full revolution.*

- 15b. Deduce for the probe in orbit that its

[4 marks]

(i) speed is  $v = \sqrt{\frac{GM}{r}}$ .

(ii) total energy is  $E = -\frac{GMm}{2r}$ .

## Markscheme

(i) equating gravitational force  $\frac{GMm}{r^2}$ ;  
to centripetal force  $\frac{mv^2}{r}$  to get result;

(ii) kinetic energy is  $\frac{GMm}{2r}$ ;  
addition to potential energy  $-\frac{GMm}{r}$  to get result;

- 15c. It is now required to place the probe in another circular orbit further away from the planet. To do this, the probe's engines will be fired for a very short time.

[2 marks]

State and explain whether the work done on the probe by the engines is positive, negative or zero.

## Markscheme

the total energy (at the new orbit) will be greater than before/is less negative;  
hence probe engines must be fired to produce force in the direction of motion / positive work must be done (on the probe);  
*Award [1] for mention of only potential energy increasing.*

### Part 2 Orbital motion

- 16a. A satellite, of mass  $m$ , is in orbit about Earth at a distance  $r$  from the centre of Earth. Deduce that the kinetic energy  $E_K$  of the satellite is equal to half the magnitude of the potential energy  $E_P$  of the satellite.

[3 marks]

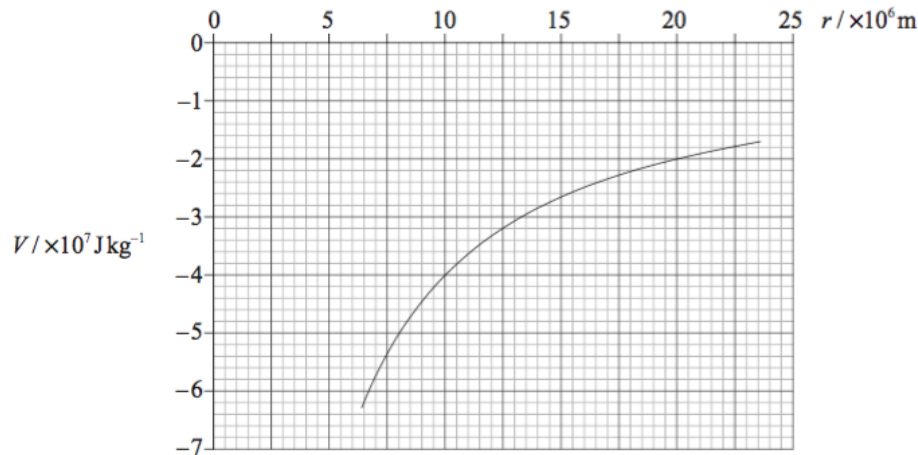
## Markscheme

$$\frac{mv^2}{r} = \frac{GMm}{r^2};$$

$$E_K = \frac{1}{2}mv^2 = \frac{GMm}{2r};$$

$$E_P = -\frac{GMm}{r} \text{ (hence magnitude of } E_K = \frac{1}{2} \text{ magnitude of } E_P);$$

- 16b. The graph shows the variation with distance  $r$  of the Earth's gravitational potential  $V$ . Values of  $V$  for  $r < R$ , where  $R$  is the radius of Earth, are not shown. [6 marks]



The satellite in (a) has a mass of  $8.2 \times 10^2 \text{ kg}$  and it is in orbit at a distance of  $1.0 \times 10^7 \text{ m}$  from the centre of Earth. Using data from the graph and your answer to (a), calculate for the satellite

- its total energy.
- its orbital speed.
- the energy it must gain to move to an orbit a distance  $2.0 \times 10^7 \text{ m}$  from the centre of the Earth.

## Markscheme

$$\begin{aligned} \text{(i) total energy} &= (\text{KE} + \text{PE}) = -\frac{Vm}{2}; \\ &= \left( -\frac{4.0 \times 10^7 \times 8.2 \times 10^2}{2} \right) = -1.6 \times 10^{10} \text{ J}; \end{aligned}$$

$$\begin{aligned} \text{(ii) } v &= \sqrt{V}; \text{ (or use of } E_K = \frac{1}{2}mv^2) \\ &= 6.3 \times 10^3 \text{ ms}^{-1}; \end{aligned}$$

$$\begin{aligned} \text{(iii) total energy in new orbit} &= \left( -\frac{2.0 \times 10^7 \times 8.2 \times 10^2}{2} \right) = -0.82 \times 10^{10} \text{ (J)}; \\ \text{energy required} &= (1.6 \times 10^{10} - 0.82 \times 10^{10}) = 7.8 \times 10^9 \text{ J}; \end{aligned}$$

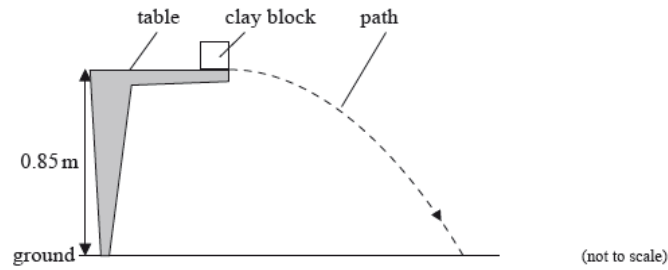
or

total energy is proportional to  $E_P$ ;  
so energy required =  $-(b)(i) \div 2 = 8$  or  $8.2 \times 10^9 \text{ J}$ ; (allow ECF from (b)(i))

This question is in **two** parts. **Part 1** is about collisions. **Part 2** is about the gravitational field of Mars.

**Part 1** Collisions

The experiment is repeated with the clay block placed at the edge of the table so that it is fired away from the table. The initial speed of the clay block is  $4.3 \text{ m s}^{-1}$  horizontally. The table surface is  $0.85 \text{ m}$  above the ground.



- 17a. (i) Ignoring air resistance, calculate the horizontal distance travelled by the clay block before it strikes the ground. [7 marks]
- (ii) The diagram in (c) shows the path of the clay block neglecting air resistance. On the diagram, draw the approximate shape of the path that the clay block will take assuming that air resistance acts on the clay block.

## Markscheme

(i) use of kinematic equation to yield time;

$$t = \sqrt{\frac{2s}{g}} (= 0.42 \text{ s});$$

$s = \text{horizontal speed} \times \text{time};$

$$= 1.8 \text{ m};$$

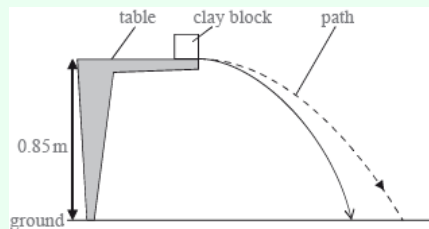
Accept  $g = 10 \text{ m s}^{-2}$

$-2$  equivalent answers 1.79 from 9.8, 1.77 from 10.

(ii) initial drawn velocity horizontal; (judge by eye)

reasonable shape; (i.e. quasi-parabolic)

horizontal distance moved always decreasing when compared to given path / range less than original;



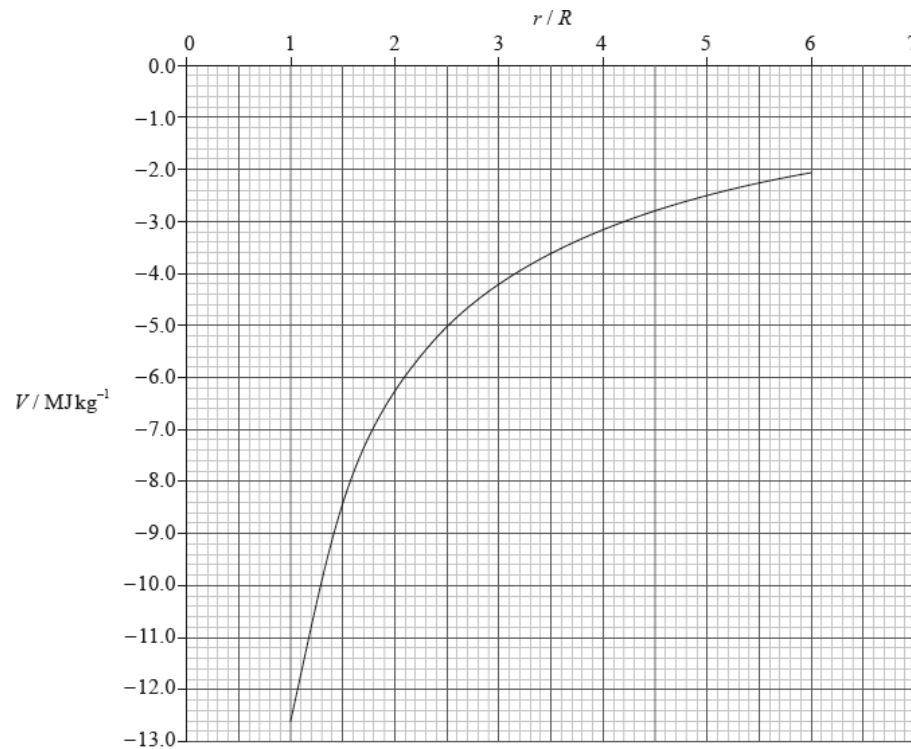
**Part 2** Gravitational field of Mars

- 17b. Define *gravitational potential energy* of a mass at a point. [1 mark]

## Markscheme

work done in moving mass from infinity to a point;

The graph shows the variation with distance  $r$  from the centre of Mars of the gravitational potential  $V$ .  $R$  is the radius of Mars which is 3.3 Mm. (Values of  $V$  for  $r < R$  are not shown.)



A rocket of mass  $1.2 \times 10^4$  kg lifts off from the surface of Mars. Use the graph to

- 17c. (i) calculate the change in gravitational potential energy of the rocket at a distance  $4R$  from the centre of Mars. [5 marks]  
 (ii) show that the magnitude of the gravitational field strength at a distance  $4R$  from the centre of Mars is  $0.23 \text{ N kg}^{-1}$ .

## Markscheme

(i) read offs  $-12.6$  and  $-3.2$ ;

gain in gpe  $1.2 \times 10^4 \times [12.6 - 3.2]$  **or** gain in g potential  $[12.6 \times 10^6 - 3.2 \times 10^6]$ ;

$= 1.13 \pm 0.05 \times 10^5 \text{ MJ}$  **or**  $1.13 \pm 0.05 \times 10^{11} \text{ J}$ ;

(ii) use of gradient of graph to determine  $g$ ;

values substituted from drawn gradient (typically  $\frac{6.7 \times 10^6}{7 \times 3.3 \times 10^6}$ );

$= 0.23 \text{ N kg}^{-1}$  (allow answers in the range of  $0.20$  to  $0.26 \text{ N kg}^{-1}$ )

Award [0] for solutions from  $\frac{V}{r}$ .